

Antiferromagnetism and its origin in iron-based superconductors

(Review Article)

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In iron-based superconductors, unravelling the origin of the antiferromagnetism is a crucial step towards understanding the high- T_c superconductivity as it is widely believed that the magnetic fluctuations play important roles in the formation of the Cooper pairs. Therefore, in this paper, we will briefly review experimental results related to the antiferromagnetic state in iron-based superconductors and focus on a review of the theoretical investigations which show applicability of the itinerant scenario to the observed antiferromagnetism and corresponding phase transitions in various families of the iron-based superconductors. A proposal of coupling between frustrated and unfrustrated bands for understanding the reduced magnetic moment typically observed in iron pnictides is also reviewed. While all the above theoretical investigations do not rule out a possible existence of localized electrons in iron-based superconductors, these results strongly indicate a close relation between itinerant electrons and the magnetically ordered state and point out the importance of taking into account the orbital degrees of freedom.

PACS: 74.70.Xa Pnictides and chalcogenides;
75.10.-b General theory and models of magnetic ordering;
71.20.-b Electron density of states and band structure of crystalline solids;
71.10.Fd Lattice fermion models (Hubbard model, etc.).

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1. Introduction

Since the first high- T_c iron-based superconductor was discovered in 2008, rapid progress has been achieved in this research field. While the measured transition temperature has been rapidly raised from 26 K [1] in LaOFeAs to 55 K as La is replaced by Sm [2], it was reported from Xue's group *et al.* [3] that there may be significantly large superconducting gap in the single-layer FeSe grown on SrTiO₃ substrate, which may open a possible avenue to synthesize new superconductors with higher transition temperature.

Apart from great experimental effort in seeking new iron-based superconductors with transition temperature higher than liquid nitrogen [4–18], much attention also has been paid to theoretical understanding of the origin of the high- T_c superconductivity [19–26]. It is now generally believed that the antiferromagnetic fluctuations must play an important role in the formation of Cooper pairs when the long range antiferromagnetic order is suppressed. This is due to the fact that the electron–phonon coupling alone cannot account for the high transition temperature observed experimentally and the superconducting phase usually appears in the vicinity of antiferromagnetic state. Therefore, it becomes a crucial starting point to first unravel the mechanism for the magnetically ordered state.

However, unlike the high- T_c cuprates where the antiferromagnetic Mott insulating state can be well interpreted from the strong coupling limit [27], intensive debates persist on whether the magnetic phases can be understood from weak coupling itinerant limit [21,28–31] or from the strong coupling localized limit [32–35] since the discovery of the first high- T_c iron-based superconductor LaFeAsO. Due to the similarities between these two high- T_c superconductors, such as the phase diagrams in the temperature–doping plane and the layered lattice structures, a common origin of the high- T_c superconductivity as well as the magnetism in both copper-based and iron-based superconductors is highly desirable. This is actually one of the reasons why the theories based on the strong coupling limit are competing with those based on the itinerant limit. Though various theoretical and experimental investigations point to a possible scenario of magnetism that exchange interactions between local spins play essential roles [32,36–40], till now, no direct evidence can confirm the existence of local spins.

On the other hand, remarkable differences are present between these two high- T_c superconductors, for example, while the undoped cuprates are antiferromagnetic Mott insulators which can be well described by a one-band Hubbard model, the parent states of iron-based high- T_c superconductors are antiferromagnetic metals with multi-bands crossing the Fermi level. As all the 3d orbitals of Fe ions give contributions to the bands crossing the Fermi level, it is evident that the itinerant electrons should be

responsible for the magnetism [21,41–43]. In fact, the itinerant point of view is supported by various theoretical and experimental results [21,28–31,44–48].

Although a compromise between these two limits was proposed that itinerant electrons and localized spins may coexist in the iron-based superconductors [49–56], more unambiguous evidences from experiments besides the angle-resolved photoemission spectroscopy study [57] are required to confirm such a coexistence. Furthermore, the mechanism for the coexistence has to be revealed beyond the slave-spin mean-field calculations [49]. After all, more sophisticated investigations like dynamical mean field theory studies did not support the localization of any electrons in 3d orbitals [58,59].

In this paper, we will first briefly review recent experimental investigations on the antiferromagnetic states in iron-based superconductors. Then, we will focus on a review of theoretical results presented by some of the authors and their collaborators and further supply new findings, which show clear evidences that magnetism in iron-based superconductors is closely related to the itinerant electrons. Our results imply that it is necessary to take into account the itinerant electrons if one would like to properly describe the physics of iron-based superconductors. However, the correlation effect can not be completely ignored as it can strongly suppress the magnetic moment, which may be responsible for the small magnetic moment observed experimentally in iron pnictides.

The remaining paper is organized as follows. In Sec. 2, experiments on the antiferromagnetic phases in different families of iron-based superconductors are presented, including effects of doping, substitution, and applying external pressure. In Sec. 3, the theoretical results based on calculations from first principles and a simplified effective model which can capture the essential physics of iron-based superconductors are shown. Conclusions are given in Sec. 4.

2. Experimental results

Up to date, a large number of iron-based superconductors have been discovered. According to the neighboring anions surrounding the irons, these iron-based materials can be cataloged into two groups, namely, iron pnictides (FePn) and iron chalcogenides (FeCh). For FePn, it can be further divided into 4 families based on their chemical formulas, including 111 (like LiFeAs and NaFeAs) [10–12], 1111 (like REOFeAs with RE = rare earth elements and AEFFeAs with AE = Ca, Sr) [1,4–6,60,61], 122 (like AEF₂As₂ where AE = Ca, Sr, Ba) [7–9] and 21311 (like Sr₂VO₃FeAs) [13]. For FeCh, there are two different structures, such as FeSe_{1-x}Te_x [14–16], called 11 family, and A_xFe_{2-y}Se₂ (A = K, Cs, Rb and Tl, etc.) [62,63,65–67], named 122* family. All the above families consist of either alternatively stacked FePn/FeCh layers and other layers or simply stacked FeCh layers. Unlike the cuprates where

only Néel ordered antiferromagnetic states are observed, different magnetic patterns have been reported in different families of the iron-based superconductors. Moreover, not only by charge-carrier doping but also by applying external pressure or isoelectronic substitution, the long range antiferromagnetic order can be suppressed and accordingly the superconductivity arises in the iron-based superconductors. Due to the complexity of the magnetic states in the iron-based superconductors, we will first briefly review the experimental results of the widely studied families.

2.1. FePn

Typically, collinear antiferromagnetic ordered states are found in most of the Fe–As compounds, except for LiFeAs [10,11] and KFe₂As₂ [64,68]. Below the spin density wave transition temperature [69], the spins on irons are arranged ferromagnetically along the orthorhombic *b* axis and antiferromagnetically along the *a* and *c* axes (note that $a > b$ here) [70]. All the Fe–P compounds are found to be nonmagnetic, such as LaOFeP [71], LiFeP [72] and BaFe₂P₂ [73].

2.1.1. 1111 structure. The magnetic moments of ReOFeAs detected in experiments are very small. Inelastic neutron scattering measurements indicate an Fe moment of $0.35\mu_B$ [74] for Re = Pr, $0.36\mu_B$ for Re = La [75], $0.25\mu_B$ for Re = Nd [76], and $0.8\mu_B$ for Re = Ce [77]. The magnetic transitions occur around 130 K, preceded by the structure phase transitions from tetragonal to orthorhombic state at about 150 K [69]. Substitution of oxygen by fluorine, served as electron doping to the Fe–As plane, can suppress both the antiferromagnetic order and structure distortion in ReOFeAs compounds [75,78,79]. By applying external pressure, it is found that the magnetic transitions are suppressed. For example, the undoped LaOFeAs exhibits a magnetic phase transition at 134 K [80] which decreases at a rate of -13.7 K/GPa as a function of applying pressure [81]. For undoped SmOFeAs, although the magnetic transition also occurs at around 130 K, similar to undoped LaOFeAs, it is found that magnetic transition decreases with pressure at a rate of -6 K/GPa [81], changing more slowly than LaOFeAs.

2.1.2. 122 structure. Inelastic neutron scattering measurements on AEF₂As₂ (AE = Ba, Ca, Sr) parent compounds show larger magnetic moments on Fe than those in the 1111 family. The magnetic moment is $0.99\mu_B$ in BaFe₂As₂ single crystal [82], $0.94\mu_B$ in SrFe₂As₂ single crystal [83] and $0.8\mu_B$ in CaFe₂As₂ [84] single crystal. Magnetic transition from nonmagnetic state to the collinear antiferromagnetic state and structure transition from *I4/mmm* tetragonal to *Fmmm* orthorhombic phase are found to occur at the same temperature in undoped AEF₂As₂ [69]. In various substitution cases such as BaFe₂(As_{1-x}P_x)₂ [85], Ba_{1-x}K_xFe₂As₂ [86,87], Ba(Fe_{1-x}Co_x)₂As₂ [88], Ba(Fe_{1-x}Ru_x)₂As₂ [89] and Ba(Fe_{1-x}Ni_x)₂As₂ [90,91], it is found that the critical temperature of magnetic and struc-

tural phase transitions decreases. Separation of these two transitions is found in some critical doping level. In the investigations of applying external pressure, it is shown that the magnetic and structure transitions are suppressed while superconductivity is induced [92]. And it is also proposed from a study of comparisons between pressure and doping effects that structural distortion or change of the Fermi surface nesting, rather than the charge carriers doping, may be responsible for the suppression of magnetic moments and appearance of superconductivity in BaFe₂As₂ [93].

2.1.3. 111 structure. LiFeAs shows superconductivity at 18 K without magnetic and structural transitions [10]. It is found from angle-resolved photoemission spectroscopy study that the Fermi surfaces are poorly nested in LiFeAs [11]. The inelastic neutron scattering and nuclear magnetic resonance studies pointed out that LiFeAs exhibits strong antiferromagnetic fluctuations [94,95]. The magnetic moment of NaFeAs is the smallest among all the iron-based superconductors which shows long range antiferromagnetic order at low temperature. It is around $(0.09 \pm 0.04)\mu_B$ from neutron scattering investigation [96] and is estimated to be $\sim 0.1-0.2 \mu_B$ from μ SR measurement [12]. Stoichiometric NaFeAs is not a bulk superconductor and can be tuned into superconducting state by small Na deficiency [96–98] or by replacing Fe by either Co or Ni [99].

2.2. FeCh

2.2.1. 11 structure. FeCh are environmentally friendly compared to the FePn compounds. The lattice structure of 11 family is simpler than that of the other families in iron-based superconductors. It consists of stacked FeCh layers only. Contrast to the Fe–As materials, Fe_{1+x}Te is found to possess a bicollinear antiferromagnetic state at $x < 0.12$ and the magnetically ordered wave vector becomes incommensurate as *x* becomes larger [100–103]. The angle-resolved photoemission spectroscopy experiments did not find the corresponding $(\pi, 0)$ nested Fermi surface topology [104]. The measured magnetic moment of Fe in Fe_{1+x}Te is about $2.25\mu_B$ [103], larger than those of Fe–As compounds. Recently it is found that Fe_{1.03}Te becomes ferromagnetic with application of pressure and the magnetic moment per Fe atom reaches about $3\mu_B$ at the highest pressure investigated [105]. Fe_{1+x}Se is nonmagnetic and shows tetragonal to orthorhombic structure transition at 90 K [106]. Superconducting transition temperature T_c of Fe_{1+x}Se gets maximum of 37 K at 7 GPa with application of hydrostatic pressure [107]. While Fe_{1+x}Te is not superconducting, the substitution of Te by Se can tune Fe_{1.02}Se_xTe_{1-x} compound from $(\pi, 0)$ magnetically ordered state to superconducting state with (π, π) magnetic resonance in the range of $0 < x < 0.5$ [108].

2.2.2. 122* structure. Alkali iron selenide superconductors A_xFe_{2-y}Se₂ with T_c around 30 K [17,65–67] are isostructural to the 122 iron-pnictide AEF₂As₂. A_xFe_{2-y}Se₂

compounds show insulating or semiconducting properties. $\text{K}_2\text{Fe}_4\text{Se}_5$ ($\text{K}_{0.8}\text{Fe}_{1.6}\text{Se}_2$) with $\sqrt{5} \times \sqrt{5}$ Fe vacancy order has the largest magnetic moment of $3.31 \mu_B$ per Fe atom among all the antiferromagnetic ordered iron-based superconductors, and the Néel temperature is up to 560 K [67]. It is demonstrated that KFe_2Se_2 with $\sqrt{2} \times \sqrt{2}$ charge ordering is the parent compound. When interfaced with $\text{K}_2\text{Fe}_4\text{Se}_5$, KFe_2Se_2 becomes superconducting with $\sqrt{2} \times \sqrt{5}$ charge ordering [109].

3. Theoretical results

From above review, several questions arise: i) Normally, the iron should be in the high spin state with $S = 2$ due to the Hund's rule coupling. So what is the origin of the reduced magnetic moment on each iron, especially in iron pnictides? ii) Contrast to the collinear antiferromagnetic state detected in iron pnictides, Fe_{1+x}Te exhibits unique bicollinear antiferromagnetic state at low temperature. So is the itinerant scenario still valid if the Fermi surface robustly shows nesting wave vector at (π, π) ? iii) Various ways such as substitution, applying pressure and charge carrier doping can suppress the antiferromagnetic phase and consequently either induce the superconductivity or become the nonmagnetic state. Can these transitions be accounted for from the itinerant picture? In this section, we will review the theoretical investigations published by some of the authors and their collaborators [29,110–115] which are related to the above questions.

3.1. Origin of reduced magnetic moment

The reduced magnetic moment observed experimentally in iron pnictides is usually ascribed to spin frustration. However, the work done by some of the authors and their collaborators [110] proposed a completely new scenario that coupling between weakly and strongly frustrated bands plays an essential role in the reduced magnetic moment. The coexistence of weakly and strongly frustrated bands is in fact a common feature in iron-based superconductors as shown in Table 1 of Ref. 116. In order to prove the validity of this scenario, a two-band Hubbard model was used which is defined as follows,

$$H = - \sum_{\langle ij \rangle, \langle\langle ij \rangle\rangle, \gamma\sigma} t_{ij,\gamma} c_{i\gamma\sigma}^\dagger c_{j\gamma\sigma} + U \sum_{i\gamma} n_{i\gamma\uparrow} n_{i\gamma\downarrow} + (U' - \frac{J}{2}) \sum_{i\gamma > \gamma'} n_{i\gamma} n_{i\gamma'} - 2J \sum_{i\gamma > \gamma'} S_{i\gamma}^z S_{i\gamma'}^z, \quad (1)$$

where $t_{ij,\gamma} = t_\gamma$ (t'_γ) is the intraorbital hopping integral between nearest-neighbor (next-nearest-neighbor) sites denoted by $\langle ij \rangle$ ($\langle\langle ij \rangle\rangle$) with orbital indices $\gamma = \alpha, \beta$ in units of t . U , U' and J are the intraband, interband Coulomb interaction and Hund's coupling, respectively, which fulfill the rotational invariance condition $U = U' + 2J$. Here $J/U = 0.25$ was chosen. The interband hybridizations, pair-hopping and spin flip terms are ignored for simplicity.

$c_{i\gamma\sigma}^\dagger$ ($c_{i\gamma\sigma}$) creates (annihilates) an electron in orbital γ of site i with spin σ . $n_{i\gamma\sigma}$ is the occupation operator, while $n_{i\gamma} = n_{i\gamma\uparrow} + n_{i\gamma\downarrow}$, and $S_{i\gamma}^z = (n_{i\gamma\uparrow} - n_{i\gamma\downarrow})/2$ the spin operator of z direction.

The model was solved by the two-sublattice dynamical mean-field theory method which includes the local quantum fluctuation effects and can account for antiferromagnetic state. The weak coupling continuous time quantum Monte Carlo method [117,118] is employed as an impurity solver. The Bethe lattice with infinite coordination is used and the self consistent conditions are given by

$$G_{0,A,\sigma}^{-1} = i\omega_n + \mu - t_\gamma^2 G_{B,\sigma} - t_\gamma'^2 G_{A,\sigma}, \quad (2)$$

$$G_{0,B,\sigma}^{-1} = i\omega_n + \mu - t_\gamma^2 G_{A,\sigma} - t_\gamma'^2 G_{B,\sigma}, \quad (3)$$

where G_0^{-1} and G are the Weiss fields and the local Green's functions, respectively. Here A and B label two different sublattices with opposite spins, ω_n is the Matsubara frequency, and μ is the chemical potential which controls the filling.

In Fig. 1, the results for the staggered magnetization per band are presented as a function of U/t for $T/t = 1/16$

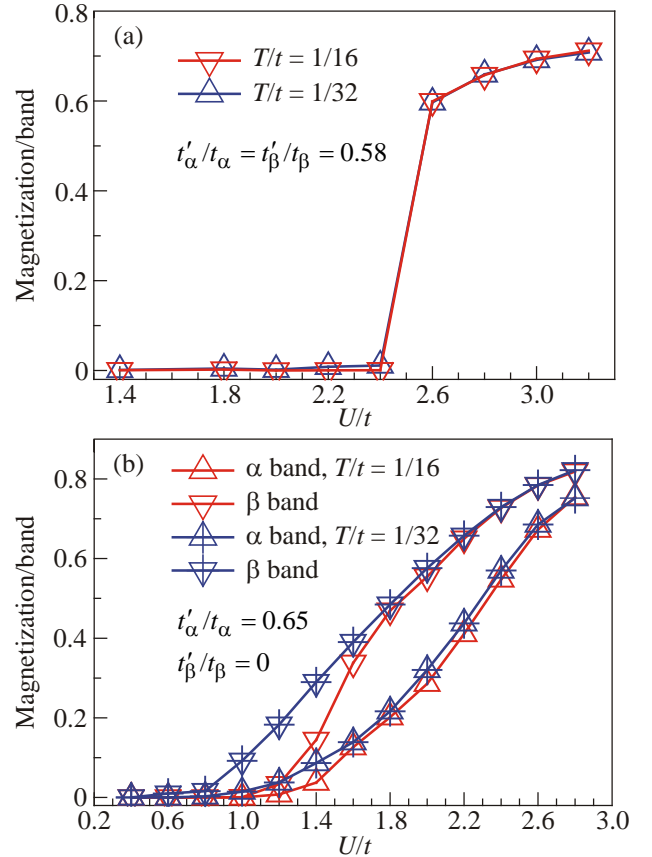


Fig. 1. (Color online) Magnetization per band of a two-band Hubbard model at two temperatures calculated by dynamical mean-field theory with weak coupling continuous time quantum Monte Carlo as a function of interaction strength U with (a) two bands equally frustrated and (b) one band frustrated, the other unfrustrated. From Ref. 110.

and 1/32. In order to find out which is the mechanism for the reduced magnetic moment, two cases were studied. One is the case of pure spin frustration as shown in Fig. 1(a). In this case, both bands are set to be equally frustrated with the ratio of $t'_\gamma/t_\gamma = 0.58$ for $\gamma = \alpha, \beta$. The other is the case of coupling between frustrated and unfrustrated bands with $t'_\alpha/t_\alpha = 0.65$ and $t'_\beta/t_\beta = 0.0$, which is shown in Fig. 1(b). Please note that the strength of the frustration is determined by the ratio of t'_γ/t_γ . Below $U/t = 2.4$ in Fig. 1(a), the staggered magnetization for both temperatures is negligibly small, indicating a paramagnetic state. As the interaction U/t is increased, for both temperatures a jump is detected around the critical value of $U_c/t = 2.4$ and the system directly evolves into an antiferromagnetic state with large magnetic moment, indicating that spin frustration alone cannot account for the state with small magnetic moment. However, in Fig. 1(b), a smooth increase is observed in the magnetization with increasing U/t for both bands and for both temperatures. While the magnetization is stronger in unfrustrated band than that in frustrated band, the magnetic moment in both bands remains small in the intermediate region of U/t , which strongly supports the mechanism of coupling between weakly and strongly frustrated bands. Certainly, local quantum fluctuations should also play important roles in stabilizing the antiferromagnetic state with small magnetic moment, as we learn from the comparisons between Hartree–Fock calculations and dynamical mean-field theory calculations [119].

3.2. Origin of bicollinear antiferromagnetism in Fe_{1+x}Te

The magnetic state of Fe_{1+x}Te observed from neutron diffraction analysis is bicollinear antiferromagnetic, which is completely different from that of iron pnictides. It requires a magnetic instability at wave vector of $(\pi, 0)$ if itinerant picture of magnetism is still valid. However, the hole and electron Fermi surfaces observed from angle-resolved photoemission spectroscopy study are still located at the center and the corner of the Brillouin zone, respectively, and separated by a wave vector of (π, π) as usually detected in iron pnictides. Such a discrepancy casts doubt on the applicability of itinerant scenario of magnetism to the 11 family and further leads to a question whether one can establish a common theory for the magnetism and superconductivity in iron pnictides and iron chalcogenides.

Recently, it was suggested by the authors [111] that magnetism in Fe_{1+x}Te still has its itinerant origin even without Fermi surface nesting, provided orbital modulation of particle–hole excitations as well as the excess Fe which donates extra electrons are considered. Magnetic exchange coupling between excess Fe and in-plane Fe further stabilizes the bicollinear antiferromagnetic order.

The scenario is verified by investigating the Pauli susceptibility, which is defined as [120]

$$\chi_{st}^{uw}(q, \omega) = -\frac{1}{N} \sum_{k, \mu\nu} \frac{a_\mu^s(k) a_\mu^{u*}(k) a_\nu^w(k+q) a_\nu^{t*}(k+q)}{\omega + E_\nu(k+q) - E_\mu(k) + i0^+} \times [f(E_\nu(k+q)) - f(E_\mu(k))], \quad (4)$$

where μ and ν are the band indices and q and k are the momentum vectors in the Brillouin zone; u, w, s, t are the orbital indices varying from 1 to 5 which represent five 3d orbitals of Fe (orbital 1: d_{xy} , 2: d_{yz} , 3: d_{xz} , 4: $d_{x^2-y^2}$, 5: d_z^2 , where x, y, z refer to those for the original unit cell). $f(E)$ is the Fermi distribution function. The matrix element $a_\mu^s(k) = \langle s | \mu k \rangle$ represents the projection of the μ -th Bloch band onto the s -th Wannier orbital at the momentum vector k , which is the component of the eigenvectors resulting from the diagonalization of an effective tight-binding Hamiltonian with ten 3d orbitals from two Fe ions and six 5p orbitals from two Te ions. The model parameters of the effective Hamiltonian are derived from the first principles calculations through construction of the maximally localized Wannier orbital basis [121,122]. The full potential linearized augmented plane wave method as implemented in Wien2k [123] is employed to calculate the band structures. The results, independent of exchange-correlation functional one chooses, are obtained within local density approximation. A three-dimensional grid of $128 \times 128 \times 128$ k and q points is used in the whole Brillouin zone with a temperature smearing of 0.01 eV for calculating the Pauli susceptibility. The experimental lattice structure [103] is used except when the substitution effect of Se by Te is studied in the system.

It was proposed that the excess Fe which can not be eliminated in the experiments [100–104] plays two roles. One is to provide extra electrons into the Fe–Te layers, which can be viewed as an electron doping. In order to take into account the doping effect, rigid band shift is used in the investigations. Figure 2 shows momentum dependent Pauli susceptibility of d_{xz} and $d_{x^2-y^2}$ orbitals along the path of $(0, 0, \pi) \rightarrow (\pi, 0, \pi) \rightarrow (\pi, \pi, \pi) \rightarrow (0, 0, \pi)$ at different Fermi energy shifts, corresponding to different number of excess Fe varying from $x = 0$ to $x = 0.2$. Here it is reasonably assumed that each excess Fe contributes 2 extra electrons into the Fe–Te layers. Without shift or as the shifts are small, like $\Delta E_F = 0$ and 0.045, both χ_{33}^{33} and χ_{44}^{44} show dominant peaks around (π, π, π) . While the peak in χ_{44}^{44} is slight enhanced at $\Delta E_F = 0.045$, compared to the case of $\Delta E_F = 0$, that in χ_{33}^{33} is strongly suppressed. At $\Delta E_F = 0.065, 0.075, 0.085$, well-defined peak centered at $(\pi, 0, \pi)$ appears in χ_{33}^{33} , indicating that bicollinear antiferromagnetic state has its itinerant origin. The peak at (π, π, π) in χ_{44}^{44} , however, remains almost unchanged at these ΔE_F . Further increasing ΔE_F to 0.105, 0.12, 0.15, the peak around (π, π, π) in χ_{44}^{44} rapidly decreases, while the peak previously right at $(\pi, 0, \pi)$ in χ_{33}^{33} moves to an incommensurate wave vector, which is consistent with experimental finding of transitions from bicollinear antiferromagnetic states to

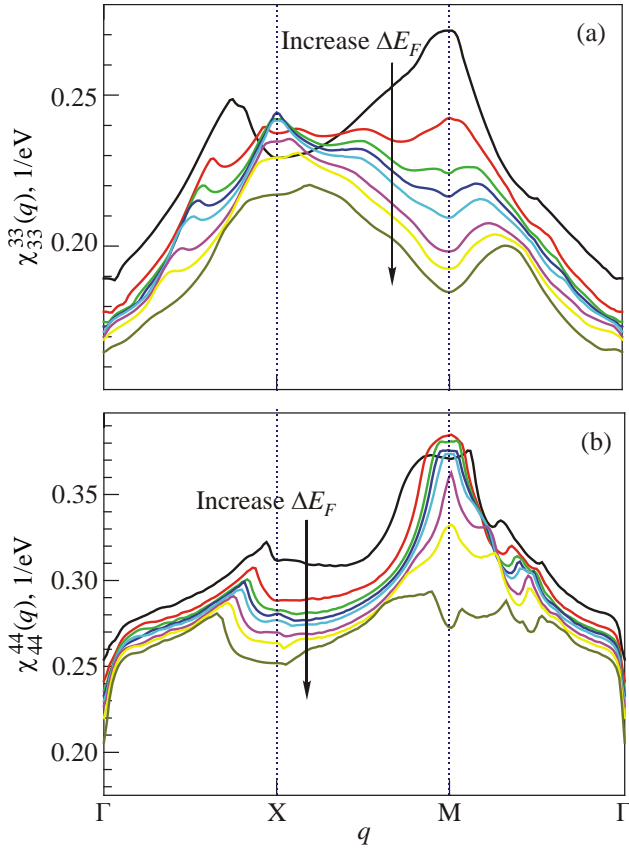


Fig. 2. (Color online) Evolution of particle-hole excitations in Pauli susceptibility as a function of shifted Fermi energies. (a) $\chi_{33}^{33}(q, \omega = 0)$ with 3 representing d_{xz} orbital. (b) $\chi_{44}^{44}(q, \omega = 0)$ with 4 denoting $d_{x^2-y^2}$ orbital. $\Delta E_F = 0, 0.045, 0.065, 0.075, 0.085, 0.105, 0.12, 0.15$ eV, corresponding to the Fe_{1+x}Te compounds with $x = 0, 0.06, 0.09, 0.10, 0.12, 0.14, 0.16, 0.20$. From Ref. 111.

incommensurate phases as number of excess Fe increases [100–103]. The larger the shift is, the farther the peak away from $(\pi, 0, \pi)$ which is also consistent with experiments. The susceptibilities of d_{z^2} and d_{xy} orbitals are negligibly small compared to those of d_{xz} and $d_{x^2-y^2}$ orbitals and that of d_{yz} orbital shows a mirror symmetry to $\chi_{33}^{33}(q, \omega = 0)$ with respect to $q_x = q_y$. Therefore, these three susceptibilities are not shown here.

Although it was shown that the bicollinear antiferromagnetic state in Fe_{1+x}Te has its itinerant origin, one may still wonder why the weaker instability of d_{xz} orbital at $(\pi, 0, \pi)$ wins the competition over the stronger instability of $d_{x^2-y^2}$ orbital at (π, π, π) . This is in fact due to the second role of the excess Fe. In addition to contributing extra electrons, the excess Fe also provides a magnetic ion which is strongly coupled with the in-plane Fe. Due to this magnetic interaction, the bicollinear antiferromagnetic state is further stabilized, rather than the collinear antiferromagnetic state. This explanation was justified by spin polarized density functional theory calculations on a supercell with 16 in-plane Fe and 1 interstitial Fe. Without

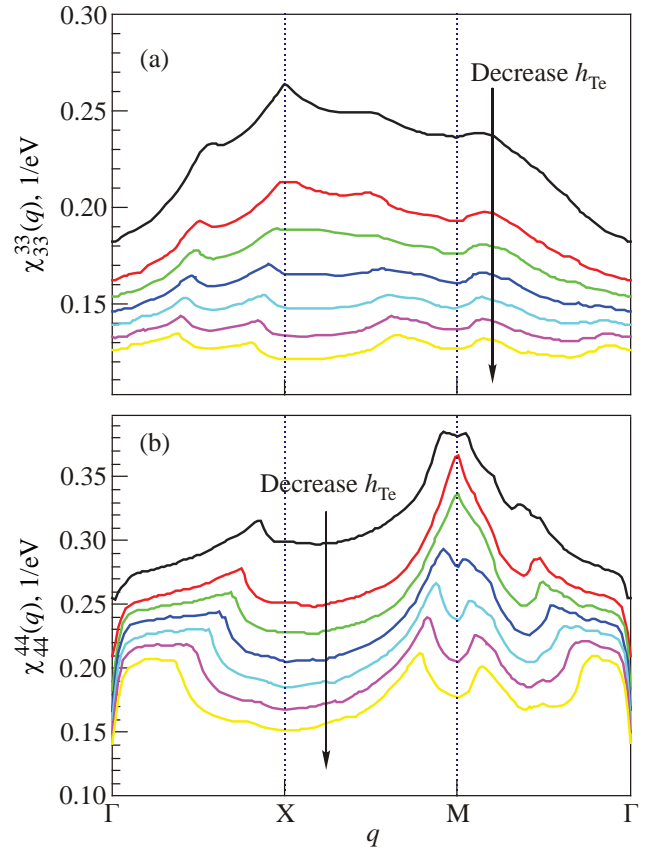


Fig. 3. (Color online) Evolution of particle-hole excitations in Pauli susceptibility as a function of Te-height measured from Fe plane. (a) $\chi_{33}^{33}(q, \omega = 0)$ with 3 representing d_{xy} orbital. (b) $\chi_{44}^{44}(q, \omega = 0)$ with 4 denoting $d_{x^2-y^2}$ orbital. Te-height from $h_{\text{Te}} = 1.8$ Å to 1.52 Å in the interval of 0.04 Å. From Ref. 111.

excess Fe, the magnetic ground state of $\text{Fe}_{16}\text{Te}_{16}$ is of collinear antiferromagnetic order within local density approximation, which is inconsistent with the experiments. After putting one excess Fe into the interstitial of the supercell with the position according to the neutron diffraction experiments, the bicollinear antiferromagnetic state turns to be the ground state, indicating that the bicollinear antiferromagnetic state is stabilized by excess magnetic Fe. However, if the magnetic moment is tuned off, which can be realized by changing the interstitial magnetic Fe to non-magnetic Zn, the collinear antiferromagnetic state becomes the ground state within local density approximation. This is a clear evidence that existence of interstitial ion with magnetic moment is crucial for stabilizing the bicollinear antiferromagnetic state.

Furthermore, evolution of the particle-hole excitations in the Pauli susceptibility as a function of Te-height measured from the Fe plane was also investigated. This can be effectively viewed as the substitution of Te by Se since Se-height is much lower than Te-height. The number of extra electrons is kept to be 0.2/Fe. From Fig. 3 one can see that

lowering Te-height rapidly suppresses the $(\pi, 0, \pi) / (0, \pi, \pi)$ instability, this may be the reason why bicollinear magnetic order disappears as substitution of Te by Se in experiment [108]. Meanwhile the (π, π, π) instability in $d_{x^2-y^2}$ orbital remaining relatively large may be the source of superconductivity with (π, π) magnetic resonance observed in experiment [108]. These results may indicate that different orbitals are responsible for different electronic properties, like magnetism or superconductivity in 11 family.

3.3. Origin of magnetic phase transitions in 122 compounds

Unlike the cuprates, various ways, like applying external pressure or charge carrier doping, etc., can be used to suppress the magnetic phase and on the other hand enhance possibly the superconducting state. In this part, we will review two typical cases which show strong evidences that the magnetic phase transitions are closely related to itinerant electrons. One is the pressure induced phase transitions occurring in AEFe_2As_2 where $\text{AE} = \text{Ca}, \text{Sr}, \text{and Ba}$. In order to capture the lattice structures under various pressures, the Car-Parrinello projector-augmented wave molecular dynamics method at constant pressure was employed [124–126]. The details can be found in Ref. 112. Here only the results for CaFe_2As_2 are shown. In Figs. 4(a) and (b), the calculated changes of the volume and magnetization of CaFe_2As_2 as a function of pressure are presented. The volume decreases gradually with increasing pressure and shows a discontinuous shrinkage at the critical pressure, where the system undergoes a structural phase transition from orthorhombic symmetry to a volume “collapsed” tetragonal symmetry. Meanwhile, the magnetization sharply goes to zero. Such a sudden jump can also be detected in the calculated Pauli susceptibility $\chi_0(q, 0)$ at $q = (\pi, \pi, 0)$ as shown Fig. 4(c). Here the $\chi_0(q, 0)$ was obtained within the constant matrix element approximation where all the $a_{\mu}^s(k)$ are set to be 1 in Eq. (4). The abrupt reduction appearing in the $\chi_0(q, 0)$ at the same critical pressure indicates that the structural and magnetic phase transitions in CaFe_2As_2 may be ascribed to the suppression of the (π, π) instability which shows tendency towards collinear antiferromagnetic order.

The second is the substitution of Fe by Co in $\text{BaFe}_{2-x}\text{Co}_x\text{As}_2$ [127]. With increase of Co concentration (x), the ground state of $\text{BaFe}_{2-x}\text{Co}_x\text{As}_2$ evolves from a collinear antiferromagnetic state ($x = 0$) to a paramagnetic state ($x = 0.5$) through an intermediate superconducting state ($x = 0.2$). Via a calculation of Pauli susceptibility $\chi_0(q, 0)$, it is also suggested that these phase transitions are closely related to the instability of the itinerant electrons. As is shown in Fig. 5 that at $x = 0$ a strong instability in the Pauli susceptibility is present at $q = (\pi, \pi, 0)$ (solid curve), supporting the collinear antiferromagnetic state. At $x = 0.2$, a strong suppression of the $q = (\pi, \pi, 0)$ instability (dashed curve) indicates a suppression of the magnetization. However, the

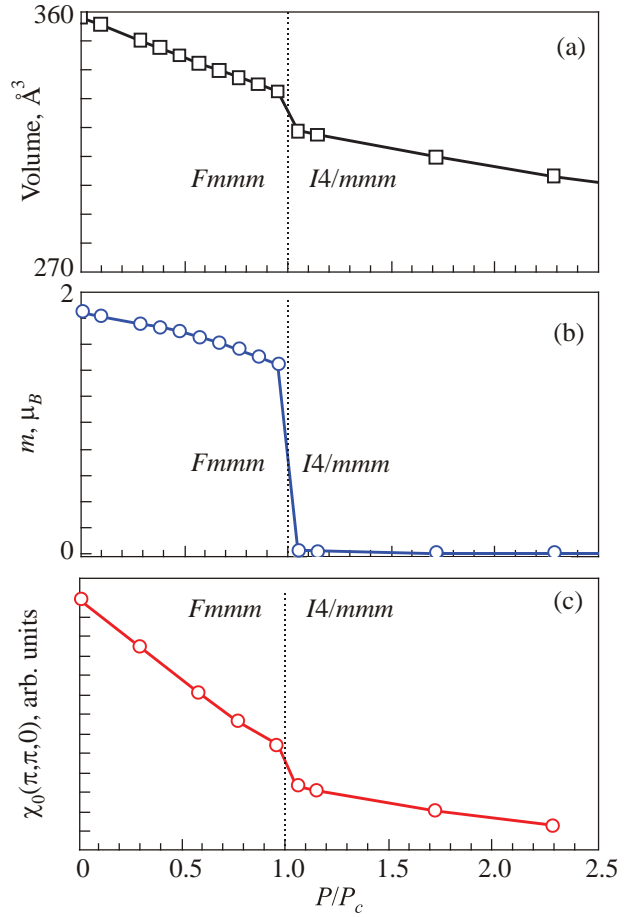


Fig. 4. (Color online) Calculated changes of volume (a), magnetization (b), and Pauli susceptibility $\chi_0(q)$ at $q = (\pi, \pi, 0)$ within the constant matrix element approximation (c) as a function of applied external pressure normalized to the critical pressure for CaFe_2As_2 . The phase boundary between $Fmmm$ and $I4/mmm$ is indicated by the vertical dashed line. From Ref. 112.

superconducting state which, according to the spin fluctuation theory [128], is related to the instabilities around $q = (\pi, \pi, 0)$ may in such a situation be more favorable than

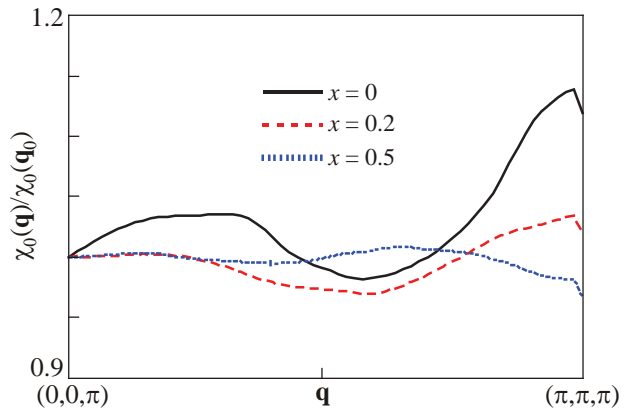


Fig. 5. (Color online) Normalized \mathbf{q} -dependent Pauli susceptibilities of $\text{BaFe}_{2-x}\text{Co}_x\text{As}_2$ at fixed $q_z = \pi$ along the [110] direction calculated within generalized gradient approximation. Here $x = 0, 0.2, 0.5$. The normalization factors are the susceptibilities of each compound at $\mathbf{q}_0 = (0, 0, \pi)$. From Ref. 113.

magnetization. At $x = 0.5$, the Pauli susceptibility $\chi_0(q, 0)$ becomes featureless and no obvious instability is present (dotted curve), leading to the disappearance of both magnetic ordering and superconductivity.

3.4. Origin of collinear antiferromagnetic states in 111 and 1111 families

In 111 family, it was found from experiments that, while NaFeAs shows collinear antiferromagnetic order at low temperature [10,96,99], it becomes nonmagnetic as Na is replaced by Li. Similar situation appears in 1111 family. While LaOFeAs exhibits collinear antiferromagnetic order, the replacement of As by P leads to a nonmagnetic state. In Ref. 29, an itinerant scenario was provided to understand the difference in the magnetic states of 1111 families. It is ascribed to a reduction of instability at $q = (\pi, \pi)$ and an increasing competition between the instabilities at $q = (\pi, \pi)$ and $q = (0, 0)$ in LaOFeP, compared to those in LaOFeAs. Here we will show that the phenomena appearing in 111 family can be also understood from the itinerant limit.

We calculate orbitally resolved Pauli susceptibilities of LiFeAs and NaFeAs according to Eq. (4). The results are presented in Fig. 6. Again we find that Pauli susceptibilities from d_{xy} and d_z^2 orbitals are negligibly small com-

pared to those from d_{xz}/d_{yz} and $d_{x^2-y^2}$ orbitals. The susceptibilities from d_{xz}/d_{yz} are almost the same in both LiFeAs and NaFeAs. However, the susceptibility from $d_{x^2-y^2}$ orbital shows remarkable difference between these two compounds. A strong peak at $q = (\pi, \pi)$ can be seen in NaFeAs, while it is absent in LiFeAs, strongly indicating that the magnetic properties of these two systems may be controlled by the $d_{x^2-y^2}$ orbital. Combining the results from 11 systems where d_{yz}/d_{xz} orbitals play important roles in the formation of magnetism and $d_{x^2-y^2}$ orbital may be responsible for the superconductivity, we propose that orbital differentiation is very important in understanding the physical properties of iron-based superconductors.

Then we will try to understand the effect of different alkali metals and the effect of changing the lattice structure. We use the experimental structures of LiFeAs [10] and NaFeAs [96]. Then we replace Li by Na in LiFeAs and replace Na by Li in NaFeAs, respectively. From Fig. 7(a) one can see that Pauli susceptibility from $d_{x^2-y^2}$ orbital becomes significantly smaller when Na is replaced by Li, implying the importance of the alkali metal to the magnetic properties. From Fig. 7(b), the Pauli susceptibility from $d_{x^2-y^2}$ orbital does not change much at the wave vector of (π, π) when Li is substituted by Na with experimental

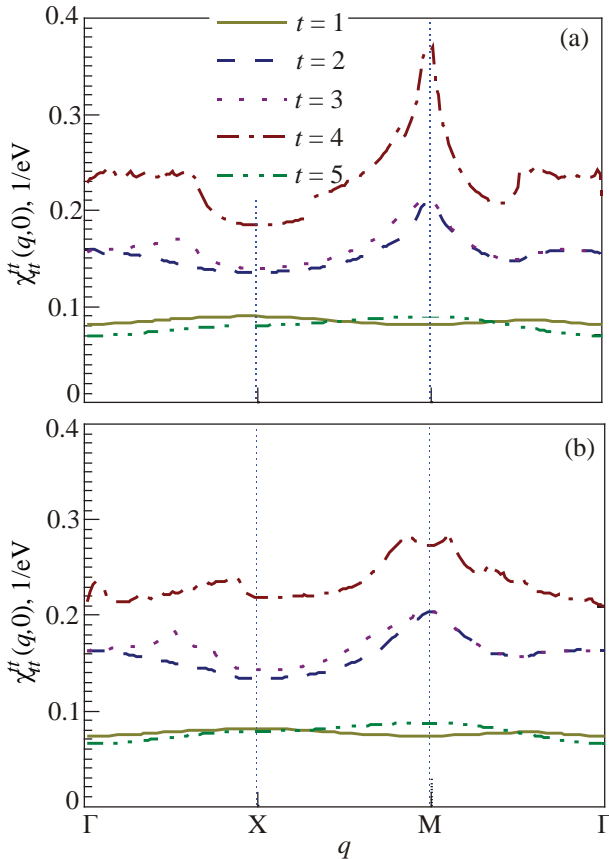


Fig. 6. (Color online) Pauli susceptibilities of NaFeAs (a) and LiFeAs (b) for all five Fe 3d orbitals. Here 1, 2, 3, 4, 5 represents d_{xy} , d_{yz} , d_{xz} , $d_{x^2-y^2}$, d_z^2 orbitals, respectively, where x, y, z refer to those for the original unit cell.

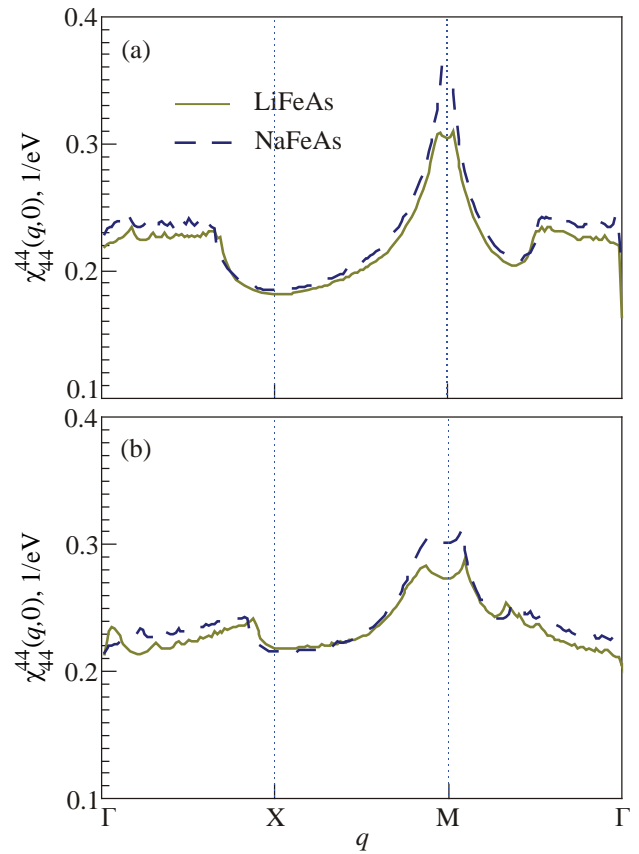


Fig. 7. (Color online) Effects of replacement of Na by Li in NaFeAs (a) and Li by Na in LiFeAs (b) on the Pauli susceptibility of $d_{x^2-y^2}$ orbital. Experimental structures of NaFeAs and LiFeAs are used in (a) and (b), respectively.

structure of LiFeAs, indicating that the lattice structure also plays an important role in determining the magnetic properties in 111 family.

4. Conclusions

In this paper, experimental progresses in the study of antiferromagnetism in iron-based superconductors are briefly reviewed. The theories from the itinerant limit established by the authors and their collaborators are also reviewed, which cover various families of iron-based superconductors, such as 122, 11, 111 and 1111. From above discussions, we can conclude that magnetism in iron-based superconductors has its itinerant origin. The inclusion of orbital degrees of freedom and the differentiation in the orbitals are very important. Strong correlations can further stabilize the small magnetic moment states. However, we do not touch the theory for the magnetism in 122* family, since many uncertainties remain from experimental sides. Most probably, the peculiar magnetic patterns observed experimentally are strongly influenced by the vacancy of the Fe. The theories presented in this paper do not rule out the possibility of existence of local spins. However, up to now, it is still lack of theory for the origin of localized electrons. Dynamical mean-field theory calculations only show heavier effective mass of quasiparticles in all the $3d$ orbitals as suitable correlation is tuned on, which casts doubt on the scenario of local spins. As dynamical mean field theory can only take into account the local fluctuations, future work should be focused on the multi-orbital physics with inclusion of spatial correlations if one would like to reveal unambiguously the origin of magnetism in iron-based superconductors.

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1. Y. Kamihara, T. Watanabe, M. Hirano, and H. Hosono, *J. Am. Chem. Soc.* **130**, 3296 (2008).
2. Z.A. Ren, W. Lu, J. Yang, W. Yi, X.L. Shen, Z.C. Li, G.C. Che, X.L. Dong, L.L. Sun, F. Zhou, and Z.X. Zhao, *Chin. Phys. Lett.* **25**, 2215 (2008).
3. Q.Y. Wang, Z. Li, W.H. Zhang, Z.C. Zhang, J.S. Zhang, W. Li, H. Ding, Y.B. Ou, P. Deng, K. Chang, J. Wen, C.L. Song, K. He, J.F. Jia, S.H. Ji, Y.Y. Wang, L.L. Wang, X. Chen, X.C. Ma, and Q.K. Xue, *Chin. Phys. Lett.* **29**, 037402 (2012).
4. C. Wang, L.J. Li, S. Chi, Z.W. Zhu, Z. Ren, Y.K. Li, Y.T. Wang, X. Lin, Y.K. Luo, S. Jiang, X.F. Xu, G.H. Cao, Z.A. Xu, *Europhys. Lett.* **83**, 67006 (2008).
5. X.H. Chen, T. Wu, G. Wu, R.H. Liu, H. Chen, and D.F. Fang, *Nature* **453**, 761 (2008).
6. G.F. Chen, Z. Li, D. Wu, G. Li, W.Z. Hu, J. Dong, P. Zheng, J.L. Luo, and N.L. Wang, *Phys. Rev. Lett.* **100**, 247002 (2008).
7. A.S. Sefat, R. Jin, M.A. McGuire, B.C. Sales, D.J. Singh, and D. Mandrus, *Phys. Rev. Lett.* **101**, 117004 (2008).
8. M.S. Torikachvili, S.L. Bud'ko, N. Ni, and P.C. Canfield, *Phys. Rev. Lett.* **101**, 057006 (2008).
9. K. Sasmal, B. Lv, B. Lorenz, A.M. Guloy, F. Chen, Y.Y. Xue, and C.W. Chu, *Phys. Rev. Lett.* **101**, 107007 (2008).
10. J.H. Tapp, Z.J. Tang, B. Lv, K. Sasmal, B. Lorenz, P.C.W. Chu, and A.M. Guloy, *Phys. Rev. B* **78**, 060505 (2008).
11. S.V. Borisenko, V.B. Zabolotnyy, D.V. Evtushinsky, T.K. Kim, Guloy, I.V. Morozov, A.N. Yaresko, A.A. Kordyuk, G. Behr, A. Vasiliev, R. Follath, and B. Büchner, *Phys. Rev. Lett.* **105**, 067002 (2010).
12. D.R. Parker, M.J. Pitcher, P.J. Baker, I. Franke, T. Lancaster, S.J. Blundell, and S.J. Clarke, *Chem. Commun.* **16**, 2189 (2009).
13. X.Y. Zhu, F. Han, G. Mu, P. Cheng, B. Shen, B. Zeng, and H.H. Wen, *Phys. Rev. B* **79**, 220512 (2009).
14. F.C. Hsu, J.Y. Luo, K.W. Yeh, T.K. Chen, T.W. Huang, P.M. Wu, Y.C. Lee, Y.L. Huang, Y.Y. Chu, D.C. Yan, and M.K. Wu, *PNAS* **105**, 38 (2008).
15. Y. Mizuguchi, F. Tomioka, S. Tsuda, T. Yamaguchi, and Y. Takano, *Appl. Phys. Lett.* **93**, 152505 (2008).
16. K.W. Yeh, T.W. Huang, Y.L. Huang, T.K. Chen, F.C. Hsu, P.M. Wu, Y.C. Lee, Y.Y. Chu, C.L. Chen, J.Y. Luo, D.C. Yan, and M.K. Wu, *Europhys. Lett.* **84**, 37002 (2008).
17. J.G. Guo, S.F. Jin, G. Wang, S.C. Wang, K.X. Zhu, T.T. Zhou, M. He, and X.L. Chen, *Phys. Rev. B* **82**, 180520(R) (2010).
18. L.L. Sun, X.J. Chen, J. Guo, P.W. Gao, Q.Z. Huang, H.D. Wang, M.H. Fang, X.L. Chen, G.F. Chen, Q. Wu, C. Zhang, D.C. Gu, X.L. Dong, L. Wang, K. Yang, A.G. Li, X. Dai, H.K. Mao, and Z.X. Zhao, *Nature* **483**, 67 (2012).
19. F. Wang and D.H. Lee, *Science* **332**, 200 (2011).
20. I.I. Mazin, *Nature* **464**, 183 (2010).
21. I.I. Mazin, D.J. Singh, M.D. Johannes, and M.H. Du, *Phys. Rev. Lett.* **101**, 057003 (2008).
22. F. Wang, H. Zhai, Y. Ran, A. Vishwanath, and D.H. Lee, *Phys. Rev. Lett.* **102**, 047005 (2009).
23. R. Thomale, C. Platt, W. Hanke, and B.A. Bernevig, *Phys. Rev. Lett.* **106**, 187003 (2011).
24. X. Lu, C. Fang, W.F. Tsai, Y. Jiang, and J. Hu, *Phys. Rev. B* **85**, 054505 (2012).
25. D.J. Scalapino, *Rev. Mod. Phys.* **84**, 1383 (2012).
26. A. Chubukov, *Annu. Rev. Condens. Matter Phys.* **3**, 57 (2012).
27. E. Dagotto, *Rev. Mod. Phys.* **66**, 763 (1994).
28. D.J. Singh, A.S. Sefat, M.A. McGuire, B.C. Sales, D. Mandrus, L.H. VanBebber, and V. Keppens, *Phys. Rev. B* **79**, 094429 (2009).

29. Y.-Z. Zhang, I. Opahle, H.O. Jeschke, and R. Valentí, *Phys. Rev. B* **81**, 094505 (2010).
30. J. Knolle, I. Eremin, A.V. Chubukov, and R. Moessner, *Phys. Rev. B* **81**, 140506 (2010).
31. D.J. Singh and M.H. Du, *Phys. Rev. Lett.* **100**, 237003 (2008).
32. B. Schmidt, M. Siahatgar, and P. Thalmeier, *Phys. Rev. B* **81**, 165101 (2010).
33. C. Fang, B. Xu, P.C. Dai, T. Xiang, and J.P. Hu, *Phys. Rev. B* **85**, 134406 (2012).
34. F.J. Ma, W. Ji, J.P. Hu, Z.Y. Lu, T. Xiang, *Phys. Rev. Lett.* **102**, 177003 (2009).
35. S. Ducatman, N.B. Perkins, and A. Chubukov, *Phys. Rev. Lett.* **109**, 157206 (2012).
36. M. Wang, C. Fang, D.X. Yao, G.T. Tan, L.W. Harriger, Y. Song, T. Netherton, C. Zhang, M. Wang, M.B. Stone, W. Tian, J. Hu, and P. Dai, *Nat. Comm.* **2**, 580 (2011).
37. J. Zhao, D.T. Adroja, D.X. Yao, R. Bewley, S. Li, X.F. Wang, G.Wu, X.H. Chen, J. Hu, and P. Dai, *Nat. Phys.* **5**, 555 (2009).
38. W. Lv, F. Krüger, and P. Phillips, *Phys. Rev. B* **82**, 045125 (2010).
39. D. Stanek, O.P. Sushkov, and G.S. Uhrig, *Phys. Rev. B* **84**, 064505 (2011).
40. J. Hu and H. Ding, *Sci. Rep.* **2**, 381 (2012).
41. J. Fink, S. Thirupathiah, R. Ovsyannikov, H.A. Dürr, R. Follath, Y. Huang, S. de Jong, M.S. Golden, Y.-Z. Zhang, H.O. Jeschke, R. Valentí, C. Felser, S.D. Farahani, M. Rotter, and D. Johrendt, *Phys. Rev. B* **79**, 155118 (2009).
42. D.J. Singh, *Phys. Rev. B* **78**, 094511 (2008).
43. A. Subedi, L. Zhang, D.J. Singh, and M.H. Du, *Phys. Rev. B* **78**, 134514 (2008).
44. S.O. Diallo, V.P. Antropov, T.G. Perring, C. Broholm, J.J. Pulikkotil, N. Ni, S.L. Bud'ko, P.C. Canfield, A. Kreyssig, A.I. Goldman, and R.J. McQueeney, *Phys. Rev. Lett.* **102**, 187206 (2009).
45. D.K. Pratt, M.G. Kim, A. Kreyssig, Y.B. Lee, G.S. Tucker, A. Thaler, W. Tian, J.L. Zarestky, S.L. Bud'ko, P.C. Canfield, B.N. Harmon, A.I. Goldman, and R.J. McQueeney, *Phys. Rev. Lett.* **106**, 257001 (2011).
46. F. Bondino, E. Magnano, M. Malvestuto, F. Parmigiani, M.A. McGuire, A.S. Sefat, B.C. Sales, R. Jin, D. Mandrus, E.W. Plummer, D.J. Singh, and N. Mannella, *Phys. Rev. Lett.* **101**, 267001 (2008).
47. Z.G. Chen, T. Dong, R.H. Ruan, B.F. Hu, B. Cheng, W.Z. Hu, P. Zheng, Z. Fang, X. Dai, and N.L. Wang, *Phys. Rev. Lett.* **105**, 097003 (2010).
48. A.A. Kordyuk, *Fiz. Nizk. Temp.* **38**, 1119 (2012) [*Low Temp. Phys.* **38**, 888 (2012)].
49. R. Yu and Q.M. Si, *Phys. Rev. Lett.* **110**, 146402 (2013).
50. Y.Z. You and Z.Y. Weng, *arXiv:1306.3740v1* (2013).
51. S.L. Yu, J. Guo, and J.X. Li, *arXiv:1305.1090v1* (2013).
52. W.G. Yin, C.C. Lee, and W. Ku, *Phys. Rev. Lett.* **105**, 107004 (2010).
53. S.-P. Kou, T. Li, and Z.-Y. Weng, *Europhys. Lett.* **88**, 17010 (2009).
54. Y.-Z. You, F. Yang, S.-P. Kou, and Z.-Y. Weng, *Phys. Rev. B* **84**, 054527 (2011).
55. Y.-Z. You, F. Yang, S.-P. Kou, and Z.-Y. Weng, *Phys. Rev. Lett.* **107**, 167001 (2011).
56. A. Hackl and M. Vojta, *New J. Phys.* **11**, 055064 (2009).
57. M. Yi, D.H. Lu, R. Yu, S.C. Riggs, J.H. Chu, B. Lv, Z.K. Liu, M. Lu, Y.T. Cui, M. Hashimoto, S.K. Mo, Z. Hussain, C.W. Chu, I.R. Fisher, Q. Si, and Z.X. Shen, *Phys. Rev. Lett.* **110**, 067003 (2013).
58. Z.P. Yin, K. Haule, and G. Kotliar, *Nature Materials* **10**, 932 (2011).
59. J. Ferber, K. Foyevtsova, R. Valentí, and H.O. Jeschke, *Phys. Rev. B* **85**, 094505 (2012).
60. Y. Xiao, Y. Su, R. Mittal, T. Chatterji, T. Hansen, C.M.N. Kumar, S. Matsuishi, Y. Inoue, H. Hosono, and Th. Brueckel, *Phys. Rev. B* **79**, 060504 (2009).
61. Y. Xiao, Y. Su, R. Mittal, T. Chatterji, T. Hansen, S. Price, C.M.N. Kumar, J. Persson, S. Matsuishi, Y. Inoue, H. Hosono, Th. Brueckel, *Phys. Rev. B* **81**, 094523 (2010).
62. Y. Zhang, L.X. Yang, M. Xu, Z.R. Ye, F. Chen, C. He, H.C. Xu, J. Jiang, B.P. Xie, J.J. Ying, X.F. Wang, X.H. Chen, J.P. Hu, M. Matsunami, S. Kimura, and D.L. Feng, *Nature Materials* **10**, 273 (2011).
63. W. Li, H. Ding, P. Deng, K. Chang, C.L. Song, K. He, L.L. Wang, X.C. Ma, J.P. Hu, X. Chen, and Q.K. Xue, *Nature Physics* **8**, 126 (2012).
64. S. Avci, O. Chmaissem, D.Y. Chung, S. Rosenkranz, E.A. Goremychkin, J.P. Castellan, I.S. Todorov, J.A. Schlueter, H. Claus, A. Daoud-Aladine, D.D. Khalyavin, M.G. Kanatzidis, and R. Osborn, *Phys. Rev. B* **85**, 184507 (2012).
65. A.K. Maziopa, Z. Shermadini, E. Pomjakushina, V. Pomjakushin, M. Bendele, A. Amato, R. Khasanov, H. Luetkens, and K. Conder, *J. Phys.: Condens. Matter* **23**, 052203 (2011).
66. M.H. Fang, H.D. Wang, C.H. Dong, Z.J. Li, C.M. Feng, J. Chen, and H.Q. Yuan, *Europhys. Lett.* **94**, 27009 (2011).
67. F. Ye, S. Chi, W. Bao, X.F. Wang, J.J. Ying, X.H. Chen, H.D. Wang, C.H. Dong, and M.H. Fang, *Phys. Rev. Lett.* **107**, 137003 (2011).
68. C.H. Lee, K. Kihou, H.K. Furukawa, T. Saito, A. Iyo, H. Eisaki, H. Fukazawa, Y. Kohori, K. Suzuki, H. Usui, K. Kuroki, and K. Yamada, *Phys. Rev. Lett.* **106**, 067003 (2011).
69. G.R. Stewart, *Rev. Mod. Phys.* **83**, 1589 (2011).
70. A.S. Sefat, *Rep. Prog. Phys.* **74**, 124502 (2011).
71. T.M. McQueen, M. Regulacio, A.J. Williams, Q. Huang, J.W. Lynn, Y.S. Hor, D.V. West, M.A. Green, and R.J. Cava, *Phys. Rev. B* **78**, 024521 (2008).
72. Z. Deng, X.C. Wang, Q.Q. Liu, S.J. Zhang, Y.X. Lv, J.L. Zhu, R.C. Yu, and C.Q. Jin, *Europhys. Lett.* **87**, 37004 (2009).
73. H. Shishido, A.F. Bangura, A.I. Coldea, S. Tonegawa, K. Hashimoto, S. Kasahara, P.M.C. Rourke, H. Ikeda, T. Terashima, R. Settai, Y. Ōnuki, D. Vignolles, C. Proust, B. Vignolle, A. McCollam, Y. Matsuda, T. Shibauchi, and A. Carrington, *Phys. Rev. Lett.* **104**, 057008 (2010).
74. S.A.J. Kimber, D.N. Argyriou, F. Yokaichiya, K. Habicht, S. Gerischer, T. Hansen, T. Chatterji, R. Klingeler, C. Hess,

- G. Behr, A. Kondrat, and B. Büchner, *Phys. Rev. B* **78**, 140503 (2008).
75. C. de la Cruz, Q. Huang, J.W. Lynn, J.Y. Li, W. Ratcliff, J.L. Zarestky, H.A. Mook, G.F. Chen, J.L. Luo, N.L. Wang, and P.C. Dai, *Nature* **453**, 07057 (2008).
76. Y. Chen, J.W. Lynn, J. Li, G. Li, G.F. Chen, J.L. Luo, N.L. Wang, P.C. Dai, C. de la Cruz, and H.A. Mook, *Phys. Rev. B* **78**, 064515 (2008).
77. J. Zhao, Q. Huang, C. de la Cruz, S.L. Li, J.W. Lynn, Y. Chen, M.A. Green, G.F. Chen, G. Li, Z. Li, J.L. Luo, N.L. Wang, and P.C. Dai, *Nature Materials* **7**, 953 (2008).
78. Y. Qiu, W. Bao, Q. Huang, T. Yildirim, J.M. Simmons, M.A. Green, J.W. Lynn, Y.C. Gasparovic, J. Li, T. Wu, G. Wu, and X.H. Chen, *Phys. Rev. Lett.* **101**, 257002 (2008).
79. C.R. Rotundu, D.T. Keane, B. Freelon, S.D. Wilson, A. Kim, P.N. Valdivia, E. Bourret-Courchesne, and R.J. Birgeneau, *Phys. Rev. B* **80**, 144517 (2009).
80. H. Luetkens, H.H. Klauss, M. Kraken, F.J. Litterst, T. Dellmann, R. Klingeler, C. Hess, R. Khasanov, A. Amato, C. Baines, M. Kosmala, O.J. Schumann, M. Braden, J. Hamann-Borrero, N. Leps, A. Kondrat, G. Behr, J. Werner, and B. Büchner, *Nature Materials* **8**, 305 (2009).
81. H. Takahashi, H. Okada, K. Igawa, Y. Kamihara, M. Hirano, H. Hosono, K. Matsubayashi, and Y. Uwatoko, *J. Supercond. Nov. Magn.* **22**, 595 (2009).
82. Y. Su, P. Link, A. Schneidewind, Th. Wolf, P. Adelman, Y. Xiao, M. Meven, R. Mittal, M. Rotter, D. Johrendt, Th. Brueckel, and M. Loewenhaupt, *Phys. Rev. B* **79**, 064504 (2009).
83. J. Zhao, W. Ratcliff, J.W. Lynn, G.F. Chen, J.L. Luo, N.L. Wang, J.P. Hu, and P.C. Dai, *Phys. Rev. B* **78**, 140504 (2008).
84. A.I. Goldman, D.N. Argyriou, B. Ouladdiaf, T. Chatterji, A. Kreyssig, S. Nandi, N. Ni, S.L. Bud'ko, P.C. Canfield, and R.J. McQueeney, *Phys. Rev. B* **78**, 100506 (2008).
85. S. Kasahara, T. Shibauchi, K. Hashimoto, K. Ikada, S. Tonegawa, R. Okazaki, H. Shishido, H. Ikeda, H. Takeya, K. Hirata, T. Terashima, and Y. Matsuda, *Phys. Rev. B* **81**, 184519 (2010).
86. D. Johrendt and R. Pöttgen, *Physica C* **469**, 332 (2009).
87. H. Chen, Y. Ren, Y. Qiu, W. Bao, R.H. Liu, G. Wu, T. Wu, Y.L. Xie, X.F. Wang, Q. Huang, and X.H. Chen, *Europhys. Lett.* **85**, 17006 (2009).
88. S. Nandi, M.G. Kim, A. Kreyssig, R.M. Fernandes, D.K. Pratt, A. Thaler, N. Ni, S.L. Bud'ko, P.C. Canfield, J. Schmalian, R.J. McQueeney, and A.I. Goldman, *Phys. Rev. Lett.* **104**, 057006 (2010).
89. M.G. Kim, D.K. Pratt, G.E. Rustan, W. Tian, J.L. Zarestky, A. Thaler, S.L. Bud'ko, P.C. Canfield, R.J. McQueeney, A. Kreyssig, and A.I. Goldman, *Phys. Rev. B* **83**, 054514 (2011).
90. P.C. Canfield, S.L. Bud'ko, Ni Ni, J.Q. Yan, and A. Kracher, *Phys. Rev. B* **80**, 060501 (2009).
91. X.Y. Lu, H. Gretarsson, R. Zhang, X.R. Liu, H.Q. Luo, W. Tian, M. Laver, Z. Yamani, Y.J. Kim, A.H. Nevidomskyy, Q.M. Si, and P.C. Dai, *Phys. Rev. Lett.* **110**, 257001 (2013).
92. A. Mani, N. Ghosh, S. Paulraj, A. Bharathi, and C.S. Sundar, *Europhys. Lett.* **87**, 17004 (2009).
93. S.A.J. Kimber, A. Kreyssig, Y.-Z. Zhang, H.O. Jeschke, R. Valentí, F. Yokaichiya, E. Colombier, J.Q. Yan, T.C. Hansen, T. Chatterji, R.J. McQueeney, P.C. Canfield, A.I. Goldman, and D.N. Argyriou, *Nature Materials* **8**, 471 (2009).
94. A.E. Taylor, M.J. Pitcher, R.A. Ewings, T.G. Perring, S.J. Clarke, A.T. Boothroyd, *Phys. Rev. B* **83**, 220514 (2011).
95. L. Ma, J. Zhang, G.F. Chen, and W.Q. Yu, *Phys. Rev. B* **82**, 180501 (2010).
96. S.L. Li, C. de la Cruz, Q. Huang, G.F. Chen, T.L. Xia, J.L. Luo, N.L. Wang, and P.C. Dai, *Phys. Rev. B* **80**, 020504 (2009).
97. K. Sasmal, B. Lv, Z.J. Tang, F. Chen, Y.Y. Xue, B. Lorenz, A.M. Guloy, and C.W. Chu, *Phys. Rev. B* **79**, 184516 (2009).
98. G.F. Chen, W.Z. Hu, J.L. Luo, and N.L. Wang, *Phys. Rev. Lett.* **102**, 227004 (2009).
99. D.R. Parker, M.J.P. Smith, T. Lancaster, A.J. Steele, I. Franke, P.J. Baker, F.L. Pratt, M.J. Pitcher, S.J. Blundell, and S.J. Clarke, *Phys. Rev. Lett.* **104**, 057007 (2010).
100. W. Bao, Y. Qiu, Q. Huang, M.A. Green, P. Zajdel, M.R. Fitzsimmons, M. Zhernenkov, S. Chang, M.H. Fang, B. Qian, E.K. Vehstedt, J.H. Yang, H.M. Pham, L. Spinu, and Z.Q. Mao, *Phys. Rev. Lett.* **102**, 247001 (2009).
101. S. Rößler, D. Cherian, W. Lorenz, M. Doerr, C. Koz, C. Curfs, Y. Prots, U.K. Rößler, U. Schwarz, S. Elizabeth, and S. Wirth, *Phys. Rev. B* **84**, 174506 (2011).
102. E.E. Rodriguez, C. Stock, P. Zajdel, K.L. Krycka, C.F. Majkrzak, P. Zavalij, and M.A. Green, *Phys. Rev. B* **84**, 064403 (2011).
103. S.L. Li, C. de la Cruz, Q. Huang, Y. Chen, J.W. Lynn, J.P. Hu, Y.L. Huang, F.C. Hsu, K.W. Yeh, M.K. Wu, and P.C. Dai, *Phys. Rev. B* **79**, 054503 (2009).
104. Y. Xia, D. Qian, L. Wray, D. Hsieh, G.F. Chen, J.L. Luo, N.L. Wang, and M.Z. Hasan, *Phys. Rev. Lett.* **103**, 037002 (2009).
105. M. Bendele, A. Maisuradze, B. Roessli, S.N. Gvasaliya, E. Pomjakushina, S. Weyeneth, K. Conder, H. Keller, and R. Khasanov, *Phys. Rev. B* **87**, 060409 (2013).
106. T.M. McQueen, A.J. Williams, P.W. Stephens, J. Tao, Y. Zhu, V. Ksenofontov, F. Casper, C. Felser, and R.J. Cava, *Phys. Rev. Lett.* **103**, 057002 (2009).
107. S. Margadonna, Y. Takabayashi, Y. Ohishi, Y. Mizuguchi, Y. Takano, T. Kagayama, T. Nakagawa, M. Takata, and K. Prassides, *Phys. Rev. B* **80**, 064506 (2009).
108. T.J. Liu, J. Hu, B. Qian, D. Fobes, Z.Q. Mao, W. Bao, M. Reehuis, S.A.J. Kimber, K. Prokeš, S. Matas, D.N. Argyriou, A. Hiess, A. Rotaru, H. Pham, L. Spinu, Y. Qiu, V. Thampy, A.T. Savici, J.A. Rodriguez, and C. Broholm, *Nature Materials* **9**, 716 (2010).
109. W. Li, H. Ding, Z. Li, P. Deng, K. Chang, K. He, S.H. Ji, L.L. Wang, X.C. Ma, J.P. Hu, X. Chen, and Q.K. Xue, *Phys. Rev. Lett.* **109**, 057003 (2012).
110. H. Lee, Y.-Z. Zhang, H.O. Jeschke, and R. Valentí, *Phys. Rev. B* **81**, 220506 (2010).
111. M.-C. Ding, H.-Q. Lin, and Y.-Z. Zhang, *Phys. Rev. B* **87**, 125129 (2013).

112. Y.-Z. Zhang, H.C. Kandpal, I. Opahle, H.O. Jeschke, and R. Valentí, *Phys. Rev. B* **80**, 094530 (2009).
113. Y.-Z. Zhang, H. Lee, I. Opahle, H.O. Jeschke, and R. Valentí, *J. Phys. Chem. Solids* **72**, 324 (2011).
114. I. Opahle, H.C. Kandpal, Y.-Z. Zhang, C. Gros, and R. Valent, *Phys. Rev. B* **79**, 024509 (2009).
115. J. Ferber, Y.-Z. Zhang, H.O. Jeschke, and R. Valent, *Phys. Rev. B* **82**, 165102 (2010).
116. Y. Yao, Y.-Z. Zhang, H. Lee, H.O. Jeschke, R. Valent, H.-Q. Lin, C.-Q. Wu, *Mod. Phys. Lett. B* **27**, 1330015 (2013).
117. A. Georges, G. Kotliar, W. Krauth, and M.J. Rozenberg, *Rev. Mod. Phys.* **68**, 13 (1996).
118. E. Gull, A.J. Millis, A.I. Lichtenstein, A.N. Rubtsov, M. Troyer, and P. Werner, *Rev. Mod. Phys.* **83**, 349 (2011).
119. Y.-Z. Zhang, H. Lee, H.-Q. Lin, C.-Q. Wu, H.O. Jeschke, and R. Valent, *Phys. Rev. B* **85**, 035123 (2012).
120. S. Graser, T.A. Maier, P.J. Hirschfeld, and D.J. Scalapino, *New J. Phys.* **11**, 025016 (2009).
121. A.A. Mostofi, J.R. Yates, Y.-S. Lee, I. Souza, D. Vanderbilt, and N. Marzari, *Comput. Phys. Commun.* **178**, 658 (2008).
122. J. Kunes, R. Arita, P. Wissgott, A. Toschi, H. Ikeda, and K. Held, *Comput. Phys. Commun.* **181**, 1888 (2010).
123. P. Blaha, K. Schwarz, G. Madsen, D. Kvaniscka, and J. Luitz, in WIEN2K, *An Augmented Plane Wave + Local Orbitals Program for Calculating Crystal*, K. Schwarz (ed.), Techn. University, Vienna, Austria (2001).
124. R. Car and M. Parrinello, *Phys. Rev. Lett.* **55**, 2471 (1985).
125. P.E. Blöchl, *Phys. Rev. B* **50**, 17953 (1994).
126. M. Parrinello and A. Rahman, *Phys. Rev. Lett.* **45**, 1196 (1980).
127. S. Thirupathiah, S. de Jong, R. Ovsyannikov, H.A. Dürr, A. Varykhalov, R. Follath, Y. Huang, R. Huisman, M.S. Golden, Y.-Z. Zhang, H.O. Jeschke, R. Valent, A. Erb, A. Gloskovskii, and J. Fink, *Phys. Rev. B* **81**, 104512 (2010).
128. T. Moriya, K. Ueda, *Rep. Prog. Phys.* **66**, 1299 (2003).