INVESTIGATION OF COMPONENTS OF THE BINARY SYSTEM 38 PER

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the binary system 38 Per are measured using high reso-kov et al., 1995), V380 Cyg (Lyubimkov et al., 1996) lution and high signal/noise ratio spectra. The orbital and 45 ε Per (Tarasov et al., 1995) confirms the corelements were updated on the basis of these data and early published measurements of V_r . The following values of effective temperature and surface gravity for the secondary component (star B) are found from its spectral type B2 V and the statistical relation of de Jager and Nieuwenhuijzen (1987): $T_{eff} = 21000 \pm 1500 \,\mathrm{K}$ $\lg g = 4.0 \pm 0.2$. Using these values we found the fundamental parameters for the primary component (star A): $T_{eff} = 22700 \pm 1500 \text{ K}$ and $\lg g = 3.4 \pm 0.2$. The rotational velocities of the components are measured from HeI lines: $v \sin i = 80 \pm 7 \text{ km s}^{-1}$ for A and $v \sin i = 95 \pm 10 \text{ km s}^{-1} \text{ for B}$.

The atmospheric helium abundance is determined; it is relatively high for A (He/H = 0.19 ± 0.06) and nearly normal for B (He/H = 0.11 ± 0.03). From evolutionary tracks the relative ages of the components $t/t_{MS}(A) = 0.96 \pm 0.05$ and $t/t_{MS}(B) = 0.54 \pm 0.03$ are determined, i.e. the primary is near the end of the main sequence (MS) phase and the secondary is near the middle of this phase. This investigation

ABSTRACT. Radial velocities V_r of components of together with our previously results for α Vir (Lyubimrelation between the helium abundance and the stellar age. An explanation may be connected with mixing already on the MS phase.

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Key words: Stars: binary: 38 Per

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