# ON THE HEURISTIC RULE FOR PLANETARY DISTANCE DISTRIBUTION 

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ABSTRACT. This paper presents a new heuristic rule for the planetary distance distribution in the solar system similar to the Titius-Bode rule of planetary orbit spacing. Application of this universal rule simultaneously for planets and planetary moons has been considered. Natural satellites orbiting around a central body are divided into groups of six satellites in each.

## 1. Introduction

There is a vast literature on the search of regularities of planetary and moon orbit spacing according to the TitiusBode - type relation [1-8]. The Titius-Bode Law of Planetary Distances: Its History and Theory by Michael M. Nieto from the Niels Bohr Institute at the University of Copenhagen was issued in 1972. Apparently, the TitiusBode relation expresses, to some extent, Newtonian mechanics in empirical form: each planet in the solar system is about 1.7 times further from the Sun than the next innermost planet. It was also shown that such regularities are realised in exoplanetary systems [2, 7]. The geometric series for distances follows from Newton's law; however, to perform sufficient simulation and deepen understanding of this phenomenon, it is necessary to rely on the methods of celestial mechanics and apply modern computer technologies. This study presents a new heuristic rule for the spacing of systems of different bodies in the solar system.

## 2. Rule definition

Natural satellites orbiting around a central body are divided into groups of six moons in each:

$$
\begin{equation*}
h_{m n}, \quad n=\overline{1.6} \tag{1}
\end{equation*}
$$

where $m$ - the group number; $n$ - the ordinal number of a moon within a group starting with the central body; $h_{m n}-$ the average distance between the central body and moon which equals to the radius of a sphere which has the same area as the planar figure restricted by the moon's orbit. If $a, b$ - the ellipse semi - axes, then the sought radius equals to $\sqrt{a b}$. The distances in the group are approximated with the following formulae (see Table 1). (Here $\alpha_{m}$ - the group non - dimensional parameter; $H_{m}$ the average orbital radius of the $6^{\text {th }}$ moon in the group, which is called the upper boundary of the group and $h_{m I}-$ the lower boundary of a group. The distances $h_{m n}$ within groups of moons are related as follows:

$$
\begin{equation*}
h_{m+k n}=\beta^{k} h_{m n}, \quad k=0,1,2,3 \ldots \tag{2}
\end{equation*}
$$

where $m+k$ - the number of a group).
Having the relative values entered, the previous table can be presented as follows (see Table 2). As is evident, here

$$
\begin{equation*}
a_{m n}=\frac{h_{m n}}{h_{m 6}} \text { and } \alpha_{m}=\frac{h_{m 6}}{h_{m 3}} \tag{3}
\end{equation*}
$$

Table 1

| Moon number | 1 | 2 | 3 | 4 | 5 | 6 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Distance notation | $h_{m 1}$ | $h_{m 2}$ | $h_{m 3}$ | $h_{m 4}$ | $h_{m 5}$ | $h_{m 6}$ |
| Distance formula | $\frac{H_{m}}{\alpha_{m}^{2}-1}$ | $\frac{H_{m}}{\alpha_{m}+1}$ | $\frac{H_{m}}{\alpha_{m}}$ | $\frac{H_{m}}{\alpha_{m}-1}$ | $\frac{2 \alpha_{m}^{H_{m}}}{\alpha_{m}^{2}-1}$ | $H_{m}$ |

Table 2

| Moon | $a_{1}$ | $a_{2}$ | $a_{3}$ | $a_{4}$ | $a_{5}$ | $a_{6}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Distance formula | $\frac{1}{\alpha^{2}-1}$ | $\frac{1}{\alpha+1}$ | $\frac{1}{\alpha}$ | $\frac{1}{\alpha-1}$ | $\frac{2 \alpha}{\alpha^{2}-1}$ | 1 |

Besides, it is supposed that the sequencing axiom is realised for planetary distances with any allowed values of the parameter $\alpha$ :

$$
\begin{equation*}
0<a_{m n}=\frac{h_{n}}{h_{6}}<a_{m n+1} \leq 1, \quad n=\overline{1,5} \tag{4}
\end{equation*}
$$

from which the left restriction for the parameter $\alpha$ is obtained:

$$
\begin{equation*}
1+\sqrt{2}<\alpha_{m} \tag{5}
\end{equation*}
$$

From the sequencing axiom, which states that the upper boundary of the group is less than the lower boundary of the next group,

$$
\begin{equation*}
h_{m-16}<h_{m n}, \quad n=1,2,3, \ldots 6 \tag{6}
\end{equation*}
$$

the right restriction is obtained:

$$
\begin{equation*}
\alpha<\sqrt{\beta+1} \tag{7}
\end{equation*}
$$

Let us suppose that there are two Phaetons rather than one hidden in the asteroid belt at the distances of 2.26 and 2.94 AU from the Sun, and that asteroid Chiron ( $a=13.65$ $\mathrm{AU}, e=0.382$ ) is a minor planet (or its remainder). According to the above - formulated rule we receive the following (see Table 3).

Asteroid 538P - L with the average orbital radius of 2.261763 AU , asteroid 1992DT2 with the average radius of 2.9403035 AU and asteroid 1999W140 with the average radius 2.9399417 AU (or the members of the Flora and Eos families) were selected as the fragments of the $5^{\text {th }}$ and $6^{\text {th }}$ planets within the first group.

Thus, the values $\alpha_{m} \approx 2.94, m=1,2,3 ; \quad \beta \approx 13$ rather accurately approximate the relative distances $\frac{h_{m k}}{h_{m 6}}, k=\overline{1.6}, m=\overline{1.3}$, obtained on the basis of actual data.

The majority of the Kuiper belt asteroids are in the region extending between the orbits of the last planet of the second group and the second planet of the third group. The $3^{\text {rd }}, 4^{\text {th }}, 5^{\text {th }}$ and $6^{\text {th }}$ planets of the third group among trans - Neptunian objects do not belong to families.

## 3. Description of moon systems

To make it more illustrative, it is more convenient to examine the Neptunian moon system first (see Table 4).

As is seen from Table 4, the first four moons make up a family of the first object within the first group. Positions for the $4^{\text {th }}$ and $5^{\text {th }}$ objects within the first group are empty. The second group is completely empty.

The existence of positions in the second group is determined by the values $\beta$, which should meet some additional requirements (see Formula 8). Besides, a definite rule, such as density axiom, can be set: in accordance with this axiom parameters $\boldsymbol{\alpha}$ and $\boldsymbol{\beta}$ should take on the least values given that conditions (2) and (6) are fulfilled.

In other words, when distributing moons within the first and third groups, the existence of the second group which fulfils condition (6) follows from condition (2).

Table 3

|  |  | PLANET | $\begin{gathered} h_{3}=\beta h_{2} \\ \beta=13.1 \end{gathered}$ | AVERAGE RADIUS | RELATION WITHIN A GROUP | THEORETICAL RELATION | $\beta$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 1 | Mercury |  | . 38294034 | . 130 | . 130 |  |
|  | 2 | Venus |  | . 722332359 | . 246 | . 254 |  |
|  | 3 | Earth |  | . 99993022 | . 340 | . 340 |  |
|  | 4 | Mars |  | 1.5203275 | . 5171 | . 516 |  |
|  | 5 | The Flora family |  | (2.2616567) | . 77 | . 77 |  |
|  | 6 | The Eos family |  | (2.94) | 1 | 1. |  |
| 2 | 1 | Jupiter |  | 5.208709 | . 134 | . 13 | 13.601882 |
|  | 2 | Saturn |  | 9.5300711 | . 245 | . 25 | 13.175292 |
|  | 3 | Chiron |  | 13.208832 | . 34 | . 34 | 13.209754 |
|  | 4 | Uranus |  | 19.18058 | . 494 | . 52 | 12.616084 |
|  | 5 | Neptune |  | 30.068409 | . 77 | . 77 | 13.258811 |
|  | 6 | Pluto |  | 38.855936 | 1. | 1. | 13.352555 |
| 3 | 1 | (229762) 2007 UK | 68.234088 | 68.69 | . 139 | . 13 | 13.2 |
|  | 2 | $\begin{aligned} & \hline(181902) 1999 \mathrm{RD}, \\ & (82158) 2001 \mathrm{FP} \\ & \hline \end{aligned}$ | 124.84393 | $\begin{gathered} \hline 105.564 \\ 157.7 \\ \hline \end{gathered}$ | $\begin{gathered} .21 \\ .319 \\ \hline \end{gathered}$ | . 25 | 11.08 |
|  | 3 | (148209) 2009 CR | 173.0357 | 169.644 | . 34 | . 34 | 12.85 |
|  | 4 | 2004 VN | 251.2656 | 243.365 | . 493 | . 52 | 12.69 |
|  | 5 | (90377) Sedna | 393.89616 | 374.634 | . 76 | . 77 | 12.46 |
|  | 6 | 2006 SQ | 509.01276 | 493.24 | 1 | 1. | 12.69 |

Table 4: The Neptunian moon system (Neptune's radius 24,764 km).

| No | $\begin{gathered} \text { Group } \\ \text { number } \end{gathered}$ | Object number within a group | Number family | Moon names | R, km | R/R | R/R theoretical | $\beta$ | $\mathbf{M}, \mathbf{k g}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | I | 1 | 1 | Naiad | 48227 | 0.1359272 | 0.11 |  | $1.9 \cdot 10^{17}$ |
| 2 |  |  | 2 | Thalassa | 50075 | 0.1411358 |  |  | $3.5 \cdot 10^{17}$ |
| 3 |  |  | 3 | Despina | 52526 | 0.1480439 |  |  | $2.1 \cdot 10^{18}$ |
| 4 |  |  | 4 | Galatea | 61953 | 0.1746380 |  |  | $2.1 \cdot 10^{18}$ |
| 5 |  | 2 |  | Larissa | 73548 | 0.2072942 | 0.24 |  | $4.9 \cdot 10^{18}$ |
| 6 |  | 3 |  | S/2004 N 1 | 105200 | 0.296505 | 0.32 |  |  |
| 7 |  |  |  | Proteus | 117647 | 0.3315858 |  |  | $5.0 \cdot 10^{19}$ |
|  |  | 4 |  | 163.5 |  |  | 0.46 |  |  |
|  |  | 5 |  | 248.6 |  |  | 0.70 |  |  |
| 8 |  | 6 |  | Triton | 354800 | 1.0000000 | 1. |  | $2.1 \cdot 10^{22}$ |
|  |  |  |  |  |  |  | $\alpha=3.17$ |  |  |
|  | II | 1 |  | 421-513.2 |  |  |  |  |  |
|  |  | 2 |  | 783-1060 |  |  |  |  |  |
|  |  | 3 |  | 1060-1252 |  |  |  |  |  |
|  |  | 4 |  | 1740-2108 |  |  |  |  |  |
|  |  | 5 |  | 2645-3500 |  |  |  |  |  |
|  |  | 6 |  | 3775-4240 |  |  |  |  |  |
|  |  |  |  |  |  |  |  | $\beta=10.64$ |  |
| 9 | III | 1 |  | Nereid | 4479360.7 | 0.0992068 | 0.125 | 9.6401649 | $3.1 \cdot 10^{19}$ |
|  |  | 2 |  | 11277.8 |  |  | 0.25 |  |  |
| 10 |  | 3 |  | Halimede | 14249954 | 0.315886 | 0.3(3) | 11.007844 | $9.0 \cdot 10^{16}$ |
| 11 |  | 4 | 1 | Sao | 21924105 | 0.4850029 | 0.5 |  | $6.7 \cdot 10^{16}$ |
| 12 |  |  | 2 | Laomedeia | 22433384 | 0.4972924 |  |  | $5.8 \cdot 10^{16}$ |
| 13 |  | 5 |  | Psamathe | 37243465 | 0.8255951 | 0.75 |  | $1.5 \cdot 10^{16}$ |
| 14 |  | 6 |  | Neso | 45111053 | 1.0000000 | 1. | 11.275853 | $1.7 \cdot 10^{17}$ |
|  |  |  |  |  |  |  | $\alpha=3.0$ |  |  |

Now let us examine the Saturnian moon system:
Drawing an analogy between macrocosm and microcosm, in accordance with the planetary model of the atom in which an electron strives to occupy the lowest orbit from the allowed ones, it can be assumed that a similar phenomenon can be observed in macrocosm as it was in the case of the $\boldsymbol{\beta}$ parameter selection during assignment of the second group of the Neptunian moons. It means that the allowed orbits of a central body's moons are determined on the same ground.

As can be seen, the moons of the Saturnian system are divided into three groups.

The moons from the $0^{\text {th }}$ to the $13^{\text {th }}$ form a sub - group located between the orbits of the first and second moons within the first group. This group can be called a family or a sub - group of the first moon within the first group.

One of the criteria by which the moons were assigned to this group, is the moons' sizes given in the last column of the table as it is not feasible to perform any other assignment.

It should be noted that the $9^{\text {th }}$ moon of the first family within the third group of the Saturnian system satisfies the following condition: $h_{3,1,9}=\beta h_{2,1}=16938$ (see Table 5).

Further let us consider the moon system of Jupiter. Using the same principles as before, we obtain data presented in Table 6. The distance for the first moon of the first group of the Jupiter system is determined by relation $h_{1}=\alpha_{1} h_{6}$, although it is less than the central body's radius (see Table 6).

The Uranian moon system can be described with four groups (see Table 7).

In different sources, the solar and planetary parameters vary significantly. Table 8 presents some variations of those parameters, as well as the obtained values of parameters $\alpha$ and $\beta$.

Table 5: The Saturnian moon system (Saturn's radius $60,268 \mathrm{~km}$ ).

$(\alpha \approx 3.2, \beta \approx 11.57, \alpha \beta \approx 37.024)$

Table 6: The moon system of Jupiter (Jupiter's radius 71,492 km).

| No | Diameter | Group number | Object number$\begin{array}{c}\text { within a } \\ \text { group }\end{array}$ | $\left\|\begin{array}{c} \text { Number } \\ \text { within a } \\ \text { family } \end{array}\right\|$ | Moon name | Average radius, thsd. km | R/R | R/R theoretical | $\beta$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | I | 1 |  | 20.49 |  |  | 0.113 |  |  |
|  |  |  | 2 |  | 43.89 |  |  | 0.242 |  |  |
|  |  |  | 3 |  | 57.67 |  |  | 0.318 |  |  |
|  |  |  | 4 |  | 84.70 |  |  | 0.467 |  |  |
| 1 | $\sim 40$ |  | 5 | 1 | Metis | 127.69 | 0.7040459 | 0.709 |  |  |
| 2 | $\sim 16$ |  |  | 2 | Adrastea | 128.69 | 0.7955960 |  |  |  |
| 3 | $\sim 146$ |  | 6 |  | Amalthea | 181.366 | 1.0 | 1.0 |  |  |
|  |  |  |  |  |  |  |  | $\alpha=3.14$ |  |  |
| 4 | -98 | II | 1 |  | Thebe | 221.872 | 0.1178490 | 0.15 | 10.83 |  |
| 5 | $\sim 3630$ |  | 2 |  | $\underline{\text { I }}$ | 421.7 | 0.2239892 | 0.26 | 9.6 |  |
| 6 | -3121,6 |  | 3 |  | Europa | 671.02 | 0.3564174 | 0.36 | 11.64 |  |
| 7 | $\sim 5262,4$ |  | 4 |  | Ganymede | 1070.412 | 0.5685575 | 0.56 | 12.64 |  |
|  |  |  | 5 |  | 1543.8 |  |  | 0.82 | 12.09 |  |
| 8 | $\sim 4820,6$ |  | 6 |  | Callisto | 1882.68 | 1.0 | 1.0 | 10.38 |  |
|  |  |  |  |  |  |  |  | $\alpha=2.8$ | $\beta=11.2$ |  |
|  |  | III | 1 |  | 2196-2622 | 阝 ¢2485 |  | 0.1153123 |  |  |
|  |  |  | 2 |  | 4634-5531 | 阝 $\mathbf{u} 4723$ |  | 0.243309 |  |  |
| 9 | 8 |  | 3 |  | Themisto | 7309.11 | 0.3214949 | 0.3214949 | 10.89 |  |
| 10 | 10 |  | 4 | 1 | Leda | 11108.66 |  | 0.4739336 |  |  |
| 11 | 170 |  |  | 2 | Himalia | 11385.86 | 0.5008128 |  | 10.63 |  |
| 12 | 86 |  |  | 3 | Elara | 11664.67 |  |  |  |  |
| 13 | 36 |  |  | 4 | Lysithea | 11688.92 |  |  |  |  |
| 14 |  |  |  | 5 | S/2000 J 11 | 12435.16 |  |  |  |  |
| 15 | 1 |  | 5 | 1 | $\underline{\text { S/2003 J } 12}$ | 16787.83 |  | 0.7172426 | 10.87 |  |
| 16 | 3 |  |  | 2 | Carpo | 16814.85 | 0.7396097 |  | 10.89 |  |
| 17 | 2 |  | 6 | 1 | Euporie | 19044.30 |  |  | 10.11 |  |
| 22 | 2 |  |  | 6 | Thelxinoe | 20074.99 |  |  | 10.66 |  |
| 32 | 28 |  |  | 16 | Ananke | 20787.92 |  |  | /каллисто=11.04 |  |
| 33 | 4 |  |  | 17 | Hermippe | 20898.75 |  |  | 11.10 |  |
| 34 | 4 |  |  | 18 | Thyone | 21055.91 | $\beta$ к=21086 |  | $\beta=10.86$ | $K_{26} \cdot \beta$ |
| 38 |  |  |  | 22 | S/2003 J 10 | 22027.12 |  |  | /каллисто=11.69 |  |
| 51 | 60 |  |  | 35 | Pasiphae | 22734.76 | $1.00000000-$ | 1.0 | /каллисто=12.08 |  |
| 54 | 46 |  |  | 38 | Carme | 22873.13 |  |  | /каллисто=12.15 |  |
| 64 | 38 |  |  | 48 | Sinope | 23589.52 |  |  | /каллисто=12.53 |  |
| 65 | 4 |  |  | 49 | Isonoe | 23610.92 |  |  | /каллисто=12.54 |  |
|  |  |  |  |  |  |  |  | $\alpha=3.11$ |  |  |
| 66 | 5 | IV | 1 | 1 | Megaclite | 24080.92 |  |  | 10.418 |  |
| 67 |  |  |  | 2 | $\underline{\mathrm{S} / 2003 \mathrm{~J} 2}$ | 30018.99 |  |  | 11.6318 |  |
|  |  |  | 2 |  | 52900 |  |  |  |  |  |
|  |  |  | 3 |  | 81900 |  |  |  |  |  |
|  |  |  | 4 |  | 12750 |  |  |  |  |  |
|  |  |  | 5 |  | 17300 |  |  |  |  |  |
|  |  |  | 6 |  | 254600 |  |  |  |  |  |
|  |  |  |  |  | $\beta=11.02$ |  |  | $\alpha=3.11$ |  |  |

$(\alpha \approx 3.04, \beta \approx 11.2, \alpha \beta \approx 34.05)$

Table 7: The Uranian moon system (the radius of Uranus is $24,800 \mathrm{~km}$ ).

| No | $\underset{\text { thsd. } \mathrm{km}}{\mathrm{R}}$ | Group number | Object number within a group | Number within a family | Moon name | r/r | $\beta$ | $\beta$ | Theoretical $\mathbf{r} / \mathbf{r}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 13.83 | I | 1 |  |  |  |  |  | 0.18445 |
|  | 21.22 |  | 2 |  |  |  |  |  | 0.28295 |
|  | 29.6 |  | 3 |  |  |  |  |  | 0.3946 |
| 1 | 49.751000 |  | 4 | 1 | Cordelia | 0.6510206 |  |  | 0.652 |
| 2 | 53.762629 |  |  | 2 | Ophelia | 0.7035151 |  |  |  |
| 7 | 66.097000 |  |  | 7 | Portia | 0.8649175 |  |  |  |
| 8 | 69.927000 |  | 5 |  | Rosalind | 0.9150353 |  |  | 0.935 |
| 9 | 74.800000 |  | 6 | 1 | Cupid | 0.9788013 |  |  | 1. |
| 10 | 75.255000 |  |  | 2 | Belinda | 0.9847553 |  |  |  |
| 11 | 76.420000 |  |  | 3 | Perdita | 1. $\alpha=2.5342$ |  |  |  |
|  |  |  |  |  |  |  | $\beta=6.25$ |  |  |
| 12 | 86.004000 | II | 1 |  | Puck | 0.1473864 | 6.23 |  | 0.126 |
| 13 | 97.734000 |  |  |  | Mab | 0.1674378 |  |  |  |
| 14 | 129.389950 |  | 2 |  | Miranda | 0.2217652 | 6.09 |  | 0.251 |
| 15 | 191.019930 |  | 3 |  | Ariel | 0.327335 | 6.453 |  | 0.334 |
| 16 | 266.298930 |  | 4 |  | Umbriel | 0.4563838 | 5.32 |  | 0.5 |
| 17 | 435.909790 |  | 5 |  | Titania | 0.7470 . | 6.236 |  | 0.75 |
| 18 | 583.519630 |  | 6 |  | Oberon | 1. $\alpha=2.99$ | 7.637 |  | 1. |
|  |  |  |  |  |  |  |  |  |  |
|  |  | III | 1 |  | 618-656 |  |  |  |  |
|  |  |  | 2 |  | 819-1113 |  |  |  |  |
|  |  |  | 3 |  | 1209-1228 |  |  |  |  |
|  |  |  | 4 |  | 1685-1743 |  |  |  |  |
|  |  |  | 5 |  | 2438-2758 |  |  |  |  |
|  |  |  | 6 |  | 3067-3692 |  |  |  |  |
|  |  |  |  |  |  |  |  | $\beta=6.39$ |  |
| 19 | 4254.116700 | IV | 1 |  | Francisco | 0.2139946 |  | 6.59 |  |
| 20 | 7218.710300 |  | 2 |  | Caliban | 0.3631319 |  | 7.469 |  |
| 21 | 7961.082700 |  | 3 | 1 | Stephano | 0.4 |  |  |  |
| 22 | 8410.678200 |  |  | 2 | Trinculo | 0.423 |  | 6.359 |  |
| 23 | 11297.873000 |  | 4 | 1 | Sycorax | 0.5682881 |  |  |  |
| 24 | 11316.714000 |  |  | 2 | Margaret | 0.569 |  | 6.519 |  |
| 25 | 15801.542000 |  | 5 | 1 | Prospero | 0.795 |  |  |  |
| 26 | 16239.657000 |  |  | 2 | Setebos | 0.8169 |  | 6.1 |  |
| 27 | 19879.088000 |  | 6 |  | Ferdinand | 1. . $\alpha=\mathbf{2 . 4 2 - 7 9 ~}$ |  | 5.837 |  |
|  |  |  |  |  |  | $\alpha=2.42$ |  |  |  |

$(\alpha \approx 2.65, \beta \approx 6.32, \alpha \beta \approx 16.75)$

Table 8: Dynamic parameters of the Sun and solar system planets.

| Planetary <br> names | The core <br> temperature, $\mathbf{T}$ | Volume (V), <br> cub. $\mathbf{m}$ | $\mathbf{I}_{O}$ | $\mathbf{I}_{O}{ }^{*}$ | $\boldsymbol{\alpha}$ | $\boldsymbol{\beta}$ | $\boldsymbol{\alpha} \boldsymbol{\beta}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Sun | $1.35-1.5 \cdot 10^{7}$ | $1.41 \cdot 10^{27}$ | 0.171 | 0.34 | 2.94 | 13.1 | 38.514 |
| Jupiter | $20-25 \cdot 10^{3}$ | $14.3-15.2 \cdot 10^{23}$ | 0,20 | 0.262 | 3.04 | 11.2 | 34.05 |
| Saturn | $11.7-20 \cdot 10^{3}$ | $8.27-9.23 \cdot 10^{23}$ | 0,22 | 0.227 | 3.2 | 11.57 | 37.024 |
| Uranus | $4.737-12 \cdot 10^{3}$ | $6.39-6.833 \cdot 10^{22}$ | 0,23 | 0.212 | 2.65 | 6.32 | 16.75 |
| Neptune | $7-14 \cdot 10^{3}$ | $6,254-6.58 \cdot 10^{22}$ | 0,26 | 0.2 | 3.1 | 10.64 | 32.984 |

Here $\mathrm{I}_{O}$ is the reduced moment of inertia.

Table 9: The S parameter values for the solar system giants and the Sun.

| Planetary <br> names | The core <br> temperature, $\mathbf{T}$ | Volume (V) <br> cub. $\mathbf{m}$ | $\mathbf{I}_{0}$ | $\boldsymbol{\alpha} \boldsymbol{\beta}$ | $\mathbf{S}$ | $\mathbf{S} \sim$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| Sun | $1.35 \cdot 10^{7}$ | $1,41 \cdot 10^{27}$ | 0.34 | 38.514 | $0.07977 \cdot 10^{20}$ | $0.08 \cdot 10^{20}$ |
| Jupiter | $25 \cdot 10^{3}$ | $1,43 \cdot 10^{24}$ | 0.2 | 34.05 | $0.08399 \cdot 10^{20}$ | $0.08 \cdot 10^{20}$ |
| Saturn | $12.15 \cdot 10^{3}$ | $8,27 \cdot 10^{23}$ | 0.22 | 37.024 | $0.08356 \cdot 10^{20}$ | $0.08 \cdot 10^{20}$ |
| Uranus | $2.45 \cdot 10^{3}$ | $6,833 \cdot 10^{22}$ | 0.2 | 16.75 | $0.08325 \cdot 10^{20}$ | $0.08 \cdot 10^{20}$ |
| Neptune | $1.2 \cdot 10^{3}$ | $6,254 \cdot 10^{22}$ | 0.2 | 32.984 | $0.08302 \cdot 10^{20}$ | $0.08 \cdot 10^{20}$ |

The given values of the parameter $S$ indirectly sustain the planetary spacing rule.

Having the values $\mathbf{T}, \mathbf{V}$ and $\mathbf{I}_{\boldsymbol{o}}$ selected (from Table 8), we see that the parameter $\mathbf{S}$, determined by the following formula:

$$
\begin{equation*}
S=\frac{V}{T 0} \frac{1}{I \alpha \beta}, \tag{8}
\end{equation*}
$$

takes on close values for planet - giants and the Sun.

## 4. Conclusions

Formally, $\alpha$, in the units of the $3^{\text {rd }}$ moon, is the upper boundary of the first group or the distance to the $6^{\text {th }}$ moon. Then, $\alpha \beta$ is the distance to the $6^{\text {th }}$ moon within the next group or the upper boundary of the second group.

Thus, a set of values $\alpha$ and $\beta$ can be determined from formulae (1) - (7) using two radii of the orbits of moons assigned to the given positions. Comparing these values with the values of $\mathrm{T}, \mathrm{V}$ and $\mathrm{I}_{O}$ in formula (8), the fittest parameter values can be found.

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